

The lowest order expansion of GW

(1)

$$h_{ij}^{\text{TT}}(t, \vec{x}) = \frac{1}{c} \frac{4G}{c^4} \Lambda_{ij,kl} \cdot S^{kl}(t - \frac{r}{c})$$

The indices of h^{TT} are latin because we can always find 4 x^ν that make $h_{0i} = 0$ and consequently $\partial^0 h_{00} = 0$ through the Lorentz condition. Hence, h_{00} is constant and we can choose it $= 0$ because we want a GW and not a feeble DC component of the gravitational field. The matrix S^{kl} is:

$$S^{kl} = \int d^3x T^{kl}(t - \frac{r}{c}, \vec{x})$$

We will demonstrate here that in order to calculate S^{kl} we need to know just the component T^{00} and not all the 6 independent components T^{kl} .

The demonstration is based on the conservation law of the energy-momentum tensor:

$$\partial_\mu T^{\mu\nu} = 0 \rightarrow \partial_0 T^{0\nu} + \partial_i T^{i\nu} = 0 \rightarrow \partial_0 T^{0\nu} = -\partial_i T^{i\nu}$$

Let's proceed. We observe that:

$$\begin{aligned} S^{kl} &= \int d^3x T^{kp} \delta_p^l = \int d^3x T^{kp} \partial_p x^l = \text{we integrate by parts} = \\ &= \int d^3x \partial_p (T^{kp} x^l) - \int d^3x \partial_p T^{kp} x^l = \text{divergence theorem} = \\ &= \underbrace{\int d^2x \hat{n}_p (T^{kp} x^l)}_{\text{Surface integral. } T^{kp} \text{ is zero outside the source}} + \int d^3x \partial_0 T^{k0} x^l = \frac{1}{c} \int d^3x \dot{T}^{k0} x^l = \\ &= \dot{P}^{kl} \end{aligned}$$

We see that there is a problem of symmetry. (2)

While $S^{kl} = S^{lk}$ because $T^{\mu\nu}$ is symmetric, the same thing cannot be said to P^{kl} : $P^{kl} \neq P^{lk}$

On the other hand we observe that we could have made a different choice at the beginning, i.e.:

$$\begin{aligned} S^{kl} &= \int d^3x T^{pl} \delta_p^k = \int d^3x T^{pl} \partial_p x^k = \dots = \int d^3x \partial_0 T^{0l} x^k = \\ &= \frac{1}{c} \int d^3x \dot{T}^{0l} x^k = \dot{P}^{l,k} \end{aligned}$$

Therefore, in order to preserve the symmetry explicitly, one could write S^{kl} in the following form:

$$S^{kl} = \int d^3x \frac{1}{2} (T^{pl} \delta_p^k + T^{kp} \delta_p^l) = \dots = \frac{1}{2} (\dot{P}^{k,l} + \dot{P}^{l,k})$$

We can continue to apply the same trick to P^{kl} and P^{lk} :

$$\begin{aligned} P^{kl} &= \frac{1}{c} \int d^3x T^{0k} x^l = \frac{1}{c} \int d^3x T^{0p} \delta_p^k x^l = \frac{1}{c} \int d^3x T^{0p} \partial_p x^k x^l = \\ &= \frac{1}{c} \int d^3x \underbrace{n_p [T^{0p} x^k x^l]}_{=0} - \frac{1}{c} \int d^3x \partial_p [T^{0p} x^k] x^l = \frac{1}{c} \int d^3x \partial_0 T^{00} x^k x^l + \\ &\quad - \frac{1}{c} \int d^3x T^{0p} \delta_p^k x^l = \frac{1}{c} \int d^3x \partial_0 T^{00} x^k x^l - P^{l,k} \end{aligned}$$

$$\text{So: } P^{kl} + P^{lk} = \frac{1}{2} \int d^3x \dot{T}^{00} x^k x^l \quad \text{Finally:}$$

$$S^{kl} = \frac{1}{2} (\dot{P}^{k,l} + \dot{P}^{l,k}) = \frac{1}{2c^2} \int d^3x \ddot{T}^{00} x^k x^l = \frac{1}{2} \ddot{\mathcal{M}}^{kl}$$